

#### Simulating the Response Matrix of a Microcalorimeter

Adam Brown and Katrina Koehler Rochester Symposium for Physics Students 15 April 2023

#### **The Systematic Problem**



H. Rotzinger, et al., J Low Temp Phys, 2008.

#### Discrepancy:

➤ Measurements ≠ Theory Prediction



#### Weak Nuclear Decay: Beta

Beta Decay:  $n \rightarrow p + e^- + \bar{\nu}_e$ 



# Reaction Energy of Beta Decay

Reaction Energy = Q Value = 
$$(M(^{N}X) - M(^{N}Y))c^{2}$$



# Distribution of Q Value Among Decay Products



# Daughter Energy Components







# Measuring Reaction Energy



# Beta Spectrum



# Beta Spectrum





600

• Only rest mass remains

400

• Upper Limit: <0.8 eV

#### Q Value Decay Energy

 $m_v$ 

Measured



#### Importance of Beta Spectrum Shape

- Determining neutrino rest mass
  - Upper Limit: <0.8 eV</li>
- Understanding Weak nuclear force

# Decay Energy Spectroscopy (DES)



### Low Temperature Detectors: Microcalorimeters



- Thermalize decay energy
- Sensitive thermometers use superconducting state
  - Small  $\Delta T \rightarrow$  Measurable change in R

### Physical Measurement



 $0.7 \text{ mm} \times 1.6 \text{ mm}$  absorber



#### **Energy Resolution**

•

High energy resolution → more detail in measurement → understand beta spectrum structure

H. Rotzinger, et al., J Low Temp Phys, 2008.

## Energy Escape





- Location
- Distribution
- Defects in Source
- Absorber size
- Particle type
- Absorber material
- Energy

#### EGSnrc: A Monte Carlo Approach



- Monte Carlo Method
- EGSnrc for Particle Transport
  - Simulate energy deposited in absorber



Compton Scattering



Photoelectric Absorption



Pair Production

#### EGSnrc: A Monte Carlo Approach



#### Monoenergetic Source Simulations



#### Finding the Detector Response: Escape



- 1 MeV monoenergetic beta source in  $0.6 \times 0.6 \times 0.6 \text{ mm}^3$  Au absorber
- Finer binning on histogram → details on spectrum
  - STRUCTURE
- Highlighted Peak:
  - 2% of electrons are only depositing 930 keV as opposed to 1 MeV
  - 70 keV is escaping absorber
  - Corresponds to Au X-ray fluorescence

#### Detector Response





Heat Map of Response Matrix

19

- 70 - 60 - 50

40

- 30

- 20 - 10

#### What Next?







• Monoenergetic electron sources?

#### EGSnrc Assumptions



- Perfect heat transfer
  - No heat lost to environment
  - Detector gain is constant
- No time dependence

# Beta Spectra

- Histograms of total reaction energy minus energy of neutrino
  - Neutrino has low likelihood of interacting with matter

Low Energy Neutrino:  

$$KE(\bullet^{e^{-}}) + E(\overset{NY}{\bullet}) = Q \text{ Value} - KE(v_e) = \qquad \Rightarrow \overset{More Reaction}{Energy Measured}$$
High Energy Neutrino:  

$$KE(\bullet^{e^{-}}) + E(\overset{NY}{\bullet}) = Q \text{ Value} - KE(v_e) = \qquad \Rightarrow \overset{Less Reaction}{Energy Measured}$$





Entranger griddioblette



#### Importance of Beta Spectrum Shape

x-axis: Energy of all products without neutrino (or Q Value minus energy of neutrino

y-axis: Frequency of energy occurring

Some other notes maybe

Determining neutrino rest mass Understanding Weak nuclear force

#### Color Palette

